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Revealing Superfluid–Mott-Insulator Transition in an Optical Lattice

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We study (by an exact numerical scheme) the single-particle density matrix of $\sim 10^3$ ultracold atoms in an optical lattice with a parabolic confining potential. Our simulation is directly relevant to the interpretation and further development of the recent pioneering experiment [1]. In particular, we show that restructuring of the spatial distribution of the superfluid component when a domain of Mott-insulator phase appears in the system, results in a fine structure of the particle momentum distribution. This feature may be used to locate the point of the superfluid–Mott-insulator transition.

The fascinating physics of the superfluid–insulator transition in a system of interacting bosons on a lattice has been attracting constant interest of theorists during recent years [2–9]. Lattice bosons is one of the simplest many-body problems with strong competition between the potential and kinetic energy, and a typical example of the quantum phase transition system. One of its great advantages is the possibility to study it by powerful Monte Carlo methods which nowadays allow simulations of many thousands of particles at low temperature with unprecedented accuracy (see, e.g., [9]). However, until very recently the canonical Bose-Hubbard model

$$H = -t \sum_{\langle ij \rangle} a_i^\dagger a_j + \frac{U}{2} \sum_i n_i^2 - \sum_i \mu_i n_i, \quad (1)$$

(where $a_i^\dagger$ creates a particle on the site i , $\langle ij \rangle$ stands for the nearest-neighbor sites, $n_i = a_i^\dagger a_i$, and t , U , and μ_i , are the hopping amplitude, the on-site interaction, and the on-site external field, respectively) was not particularly useful in the analysis of realistic systems. The situation has changed with the exciting success of the experiment by Greiner *et al.* [1] (originally proposed by Jaksch *et al.* [10]) in which a gas of ultracold ^{87}Rb atoms was trapped in a three-dimensional, simple-cubic optical lattice. The uniqueness of the new system is that it is adequately described by the Bose-Hubbard Hamiltonian [10,1], and allows virtually an unlimited control over the strength of the effective interparticle interaction U/t and particle density.

The characteristic feature of the experimental setup of Ref. [1] is the presence of the overall parabolic potential $V(r)$ which confines the sample. This feature could be of great advantage if one was able to directly measure the spatial density distribution in the trap. We recall the structure of the $\mu - U/t$ phase diagram for the Bose-Hubbard system [2] which predicts commensurate particle density distribution for the insulating phase whenever the chemical potential lies within the Mott-Hubbard gap. The slowly varying (at the length scale of the lattice pe-

riod) trapping potential effectively provides a scan over μ of this phase diagram at a fixed value of U/t .

Unfortunately, what is measured in the experiment is not the original spatial density distribution in the trap, but the absorption image of the free evolving atomic cloud, after the trapping/optical potential is removed; i.e. the quantity which is directly related to the single-particle density matrix in *momentum* space, $\rho_{\mathbf{k}\mathbf{k}} = n_{\mathbf{k}}$. [This statement implies that in the free evolving atomic cloud the interparticle interaction can be neglected; see the discussion below.] Now, in terms of $n_{\mathbf{k}}$ the inhomogeneous trapping potential is a disadvantage since it broadens the superfluid δ -functional contribution at $\mathbf{k} = 0$, and the observed picture is a convolution of the original real-space density matrix $\rho(\mathbf{r}, \mathbf{r}')$. As we show below one has then to look at the fine-structure of the central peak in the experimental data to decipher the Mott-Hubbard phase diagram (the “fading” of the Bragg peaks in the experiment has very little to do with the superfluid–Mott-insulator transitions and happens when the system is already deep in the insulating phase).

In this Letter we relate quantitatively the particle distribution in momentum space observed in experiments to the corresponding spatial density distribution in the trap. Our ultimate goal is to reveal which features (if any) in the structure of $n_{\mathbf{k}}$ indicate unambiguously the presence of the Mott phase. To this end we perform quantum Monte Carlo simulations of the single-particle density matrix for the Bose-Hubbard systems with up to 16^3 lattice sites using continuous-time Worm algorithm [11]. We find that the onset of the phase transition in the trap center should result in appearance of at least one satellite peak in $n_{\mathbf{k}}$, reflecting a shell-type form of the superfluid component. This peak was not mentioned in the experiments of Ref. [1]. We suggest a possible explanation to this fact, and argue that by collimating the expanding atomic cloud one can render this peak observable. We also discuss the role of self-repulsion in the free expanding cloud, which can affect the simple interpretation of the absorption images in terms of the initial

single-particle momentum distribution.

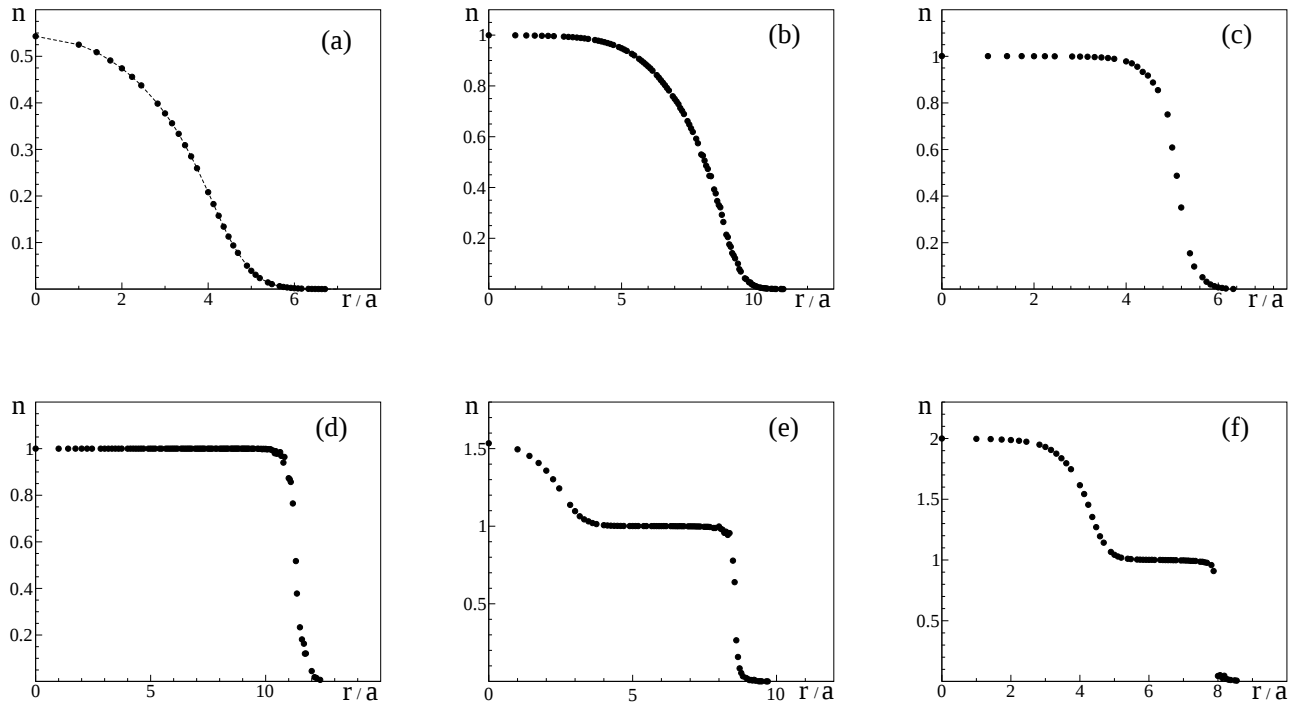


FIG. 1. Particle density distributions (on-site filling factors) as functions of the lattice site distance from the trap center for various coupling parameters and filling factors in the center: $U/t = 24$ (a), $U/t = 32$ (b), $U/t = 80$ (c), $U/t = 80$ (d), $U/t = 80$ (e), $U/t = 80$ (f).

In Figs. 1(a-f) we present our data for the density distribution as function of the lattice site distance from the trap center, r/a , where a is the lattice constant. For all practical purposes one may assume the zero-temperature limit here. The simulation was done at finite but very low $T = 0.2t$; the relevant energy parameters in this model are the bandwidth $W = 12t$ and $(U/t)_c \sim 35$ [3,4]. In accordance with the phase diagram of Ref. [2], we observe a shell-type structure of the particle density with the Mott-insulator phases visible as integer plateau regions.

Next, we relate each of the above figures to the corresponding momentum distribution function $n_{\mathbf{k}}$. By definition, $n_{\mathbf{k}} = \int d^3r d^3r' \exp[i\mathbf{k}(\mathbf{r} - \mathbf{r}')] \rho(\mathbf{r}, \mathbf{r}')$, where $\rho(\mathbf{r}, \mathbf{r}') = \langle \psi^\dagger(\mathbf{r}) \psi(\mathbf{r}') \rangle$, and $\psi(\mathbf{r})$ is the bosonic field operator. In our case of a single-zone lattice, the field operator is expanded as follows:

$$\psi(\mathbf{r}) = \sum_i \phi(\mathbf{r} - \mathbf{r}_i) a_i, \quad (2)$$

where ϕ is the Wannier function. We thus finally have

$$n_{\mathbf{k}} = |\phi(\mathbf{k})|^2 \sum_{i,j} e^{i\mathbf{k}(\mathbf{r}_i - \mathbf{r}_j)} \rho_{ij}, \quad (3)$$

where

$$\rho_{ij} = \langle a_i^\dagger a_j \rangle, \quad (4)$$

and $\phi(\mathbf{k})$ is the Fourier transform of $\phi(\mathbf{r})$. From Eq. (3) it is seen that up to a trivial reweighting factor $|\phi(\mathbf{k})|^2$ the distribution is a periodic function in the reciprocal lattice. Thus without loss of generality we may restrict ourselves to the first Brillouin zone. Actually, $\phi(\mathbf{k})$ has nothing to do with the Bose-Hubbard model, being a non-universal property of the lattice site potential; in what follows we will ignore this function altogether by formally setting it to unity.

Having calculated ρ_{ij} with the Worm algorithm [11], we readily obtain $n_{\mathbf{k}}$ using Eq. (3); the results are presented in Figs. 2(a-f). In Fig. 2(a) we see a typical picture for the strongly correlated superfluid phase, characterized by a single, narrow peak at small momenta. When a domain of the Mott-insulating phase appears

in the center of the trap (where the on-site filling is close to unity), a pronounced fine structure develops in Fig. 2(b). We associate this structure with the shell-type form of the condensate wave-function. To prove the point we model the situation with the pure-condensate density matrix $\rho(\mathbf{r}, \mathbf{r}') = \Psi_0^*(\mathbf{r})\Psi_0(\mathbf{r}')$, where the condensate wave-function $\Psi_0(\mathbf{r})$ has the shell-type form with the shell radius l . The presence of the Mott insulator is taken into account through the suppression of the $\Psi_0(\mathbf{r})$ in the center. The Fourier transform of such $\Psi_0(\mathbf{r})$ is alternating in sign, with the half-period in k related to

the shell radius as $k \sim \pi/l$. Thus in the pure condensate we would see exact zeros in $n_{\mathbf{k}}$ with the typical separation between them $\sim \pi/l$. Surprisingly, this naive model works extremely well and adequately describes the case of the realistic strongly correlated system close to the phase transition (in Fig. 2(b) the coupling $U/t = 32$ is close to the critical value estimated in Refs. [3,4]). We consider the appearance of the satellite peaks as a clear signature of the Mott-insulator transition in the center of the trap.

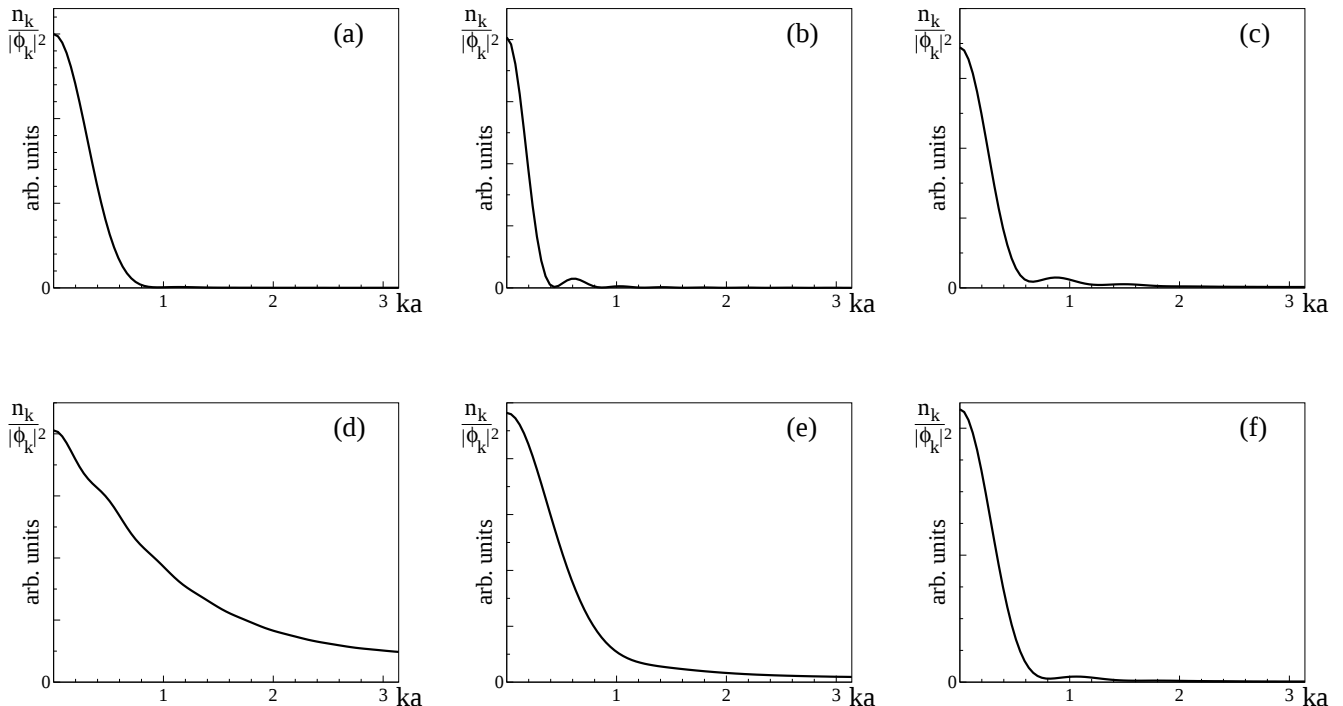


FIG. 2. The $n_{\mathbf{k}}$ distributions in the first Brillouin zone in the (0,0,1) direction derived from the single-particle density matrices for systems shown in Figs. 1(a-f).

In Fig. 2(c) the coupling strength is significantly increased to $U/t = 80$, but the shell-type structure of the superfluid phase, and the corresponding fine-structure in $n_{\mathbf{k}}$, is still present. The crucial difference between Fig. 2(c) and Fig. 2(d) is in the suppression (almost complete) of the superfluid fraction. Now the distribution $n_{\mathbf{k}}$ has only the central peak and is extended towards large momenta, as expected for the system deep in the insulating phase. Still, it is not flat (!), which tells us about large off-diagonal correlations between the nearest lattice sites even for U/t as large as $\sim 2.5 (U/t)_c$.

Fig. 2(e) is similar in physics to the case (a), except

for the large-momentum tail due to the Mott-insulator shell. Finally, in Fig. 2(f) we again see the fine structure of satellite peaks reflecting the appearance of the Mott-insulator phase in the center of the trap and the corresponding superfluid phase shell [in close resemblance with Fig. 2(b)].

We note, that momentum distributions $n_{\mathbf{k}}$ presented above, may be observed experimentally if atoms are collimated out of the expanding cloud so that the distribution in a given direction is photographed. In the current setup, the absorption images of the three dimensional distribution are taken along two orthogonal axis.

This procedure reveals only the integral $N(k_x, k_y) \propto \int_{-\infty}^{\infty} dk_z n(\mathbf{k})$. It is easy to see that integration effectively erases fine-structure features of $n(\mathbf{k})$ —although peaks do not disappear completely, they now show up as shoulders in $N(k_x, k_y)$.

Finally, we would like to discuss the problem of the repulsive interaction between particles during the initial period of their free expansion. Obviously, the interpretation of the final absorption pattern in terms of the *initial* momentum distribution is valid only if the effect of this interaction is small enough. Meanwhile, given realistic experimental parameters, this turns out to be the case only for rather moderate system sizes. The criterion for neglecting the effect of interparticle repulsion is

$$E_{\text{kin}}/E_{\text{pot}} \gg 1, \quad (5)$$

where E_{kin} and E_{pot} are, respectively, the kinetic and potential energy per particle in the most fragile low-momentum part of the distribution $n_{\mathbf{k}}$ and at the most dangerous period of free evolution at the end of the restructuring period, when the “discrete” distribution of density transforms into the “continuous” spatial distribution with the typical size of order of the original system size (plus the corresponding replicas in higher Brillouin zones). For the potential energy we have $E_{\text{pot}} \sim nU_0$, where n is the continuous number density, $U_0 = 4\pi\hbar^2 a_s/m$, a_s is the s -scattering length, and m is the atom mass. Recalling that the lattice filling factor is of order unity, we can estimate $n \sim 1/a^3$. The lowest kinetic energy is associated with the spatial distribution of the condensate. Estimating $E_{\text{kin}} \sim \pi^2\hbar^2/ma^2L^2$, where integer L stands for the typical size of the superfluid component in units of lattice constants, we arrive at a simple requirement

$$(a/a_s)L^{-2} \gg 1. \quad (6)$$

In the experiment of Ref. [1], the ratio a/a_s is of order 10^2 . Hence, we are restricted to $L < 10$, that is to typical system sizes of our present simulation. Note also, that the condition (5) is much easier to satisfy for the atomic cloud in the *second* Brillouin-zone peak, where the spatial density is significantly suppressed [1].

Summarizing, we presented a simulation of the ground state properties of ultracold atoms in an optical lattice with confining external potential, in the regime where the Mott-insulator and superfluid phases coexist. We

have demonstrated that when the insulator domain in the center of the trap is surrounded by the superfluid component, the global momentum distribution of particles features satellite peaks. This picture can be employed by the experiment as an unambiguous evidence of the Mott transition. We do not see other features of the momentum distribution that could be associated with the Mott transition: unless the ratio U/t is not much larger than the critical one, the momentum distribution in the reciprocal lattice still has a peaked form reflecting strong local off-diagonal correlations.

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