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天体物理重要反应¹³N(p,**v**)¹⁴O的实验研究

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摘要 $^{13}N(p,\gamma)^{14}O$ 是高温CNO循环中的关键反应,对恒星能量产生机制及其演化的研究具有重要意义。利用北京HI-13串列加速器次级束流线产生的13N放射性束测量了质心系能量为8.9 MeV的 $^{13}N(d,n)^{14}O$ 反应的角分布,导出了 ^{14}O 基态渐进归一化系数(ANC)为(29.4±5.3) fm $^{-1}$ 。此外,使用镜像核的电荷对称性,通过分析 $^{13}C(d,p)$ ^{14}C 反应的角分布,导出了与实验一致的 ^{14}O 基态质子ANC。使用最新开发出的 R 矩阵程序,导出 $^{13}N(p,\gamma)^{14}O$ 反应在高温CNO循环中的天体物理S因子和反应率。将此数据代入核天体物理反应的网络程序进行计算,结果表明,新星中CNO循环产生的能量比原有的结果多5%,这可能会对新星的演化有一定的影响。

关键词 <u>高温CNO循环</u> 放射性束流 角分布 渐进归一化系数 <u>S因子</u> 分类号

Experimental Study of $^{13}N(p, \mathbf{y})^{14}O$ Reaction of Astrophysical Importance

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Abstract $13N(p,\gamma)14O$ is one of the key reaction in hot CNO cycle, which is important for stud ying the energy source of evolution of stars. The angular distribution of 13N(d,n)14O at Ecm = 8.9 MeV was measured using 13N beam produced by secondary beam facility at HI-13 tande m accelerator. The ANC of 14O was then extracted to be (29.4 ± 5.3) fm-1. In addition, we derived the asymptotic normalization coefficient (ANC) of 14O by analyzing the 13C(d,p)14C angular distribution based on charge symmetry of mirror nuclei, which agrees with the experimental result. The astrophysical S factor and reaction rate of $13N(p,\gamma)14O$ were then calculated using R-mat rix theory and were inputted into astrophysical reaction network. It shows the energy generated by CNO cycle in novae using the present data is 5% more than that using the previous ones, which could affect the evolution of novae.

 Key words
 hot
 CNO
 cycle
 radioactive
 nuclear
 beam
 angular
 distribution

 asymptotic
 normalization
 coefficient
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 S
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扩展功能

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