

Astrophysics > Cosmology and Extragalactic Astrophysics

Evolution of Infrared Luminosity functions of Galaxies in the AKARI NEP-Deep field: Revealing the cosmic star formation history hidden by dust

Tomotsugu Goto (ifa, University of Hawaii), T. Takagi, H. Matsuhara (ISAS/JAXA), T.T. Takeuchi (Nagoya), C. Pearson (Rutherford), T. Wada, T. Nakagawa (ISAS/JAXA), O. Ilbert (Marseille), E. Le Floc'h (CEA), S. Oyabu (ISAS/JAXA), Y. Ohyama (ASIAA), M. Malkan (UCLA), H.M. Lee, M.G. Lee (SNU), H. Inami (SSC), N. Hwang (NAOJ), H. Hanami (Iwate), M. Im (SNU), K. Imai (TOME R&D), T. Ishigaki (Asahikawa), S. Serjeant (Open Univ.), H. Shim (SNU)

(Submitted on 30 Dec 2009 (v1), last revised 8 Jan 2010 (this version, v2))

Dust-obscured star-formation becomes much more important with increasing intensity, and increasing redshift. We aim to reveal cosmic star-formation history obscured by dust using deep infrared observation with the AKARI.

We construct restframe 8 μ m, 12 μ m, and total infrared (TIR) luminosity functions (LFs) at $0.15 < z < 2.2$ using 4128 infrared sources in the AKARI NEP-Deep field. A continuous filter coverage in the mid-IR wavelength (2.4, 3.2, 4.1, 7, 9, 11, 15, 18, and 24 μ m) by the AKARI satellite allows us to estimate restframe 8 μ m and 12 μ m luminosities without using a large extrapolation based on a SED fit, which was the largest uncertainty in previous work. We have found that all 8 μ m ($0.38 < z < 2.2$), 12 μ m ($0.15 < z < 1.16$), and TIR LFs ($0.2 < z < 1.6$), show a continuous and strong evolution toward higher redshift. In terms of cosmic infrared luminosity density (Ω_{IR}), which was obtained by integrating analytic fits to the LFs, we found a good agreement with previous work at $z < 1.2$, and that the Ω_{IR} evolves as $\propto (1+z)^{4.4-1.0}$. When we separate contributions to Ω_{IR} by LIRGs and ULIRGs, we found more IR luminous sources are increasingly more important at higher redshift. We found that the ULIRG (LIRG) contribution increases by a factor of 10 (1.8) from $z=0.35$ to $z=1.4$.

Comments: Accepted for publication in A&A AKARI special issue

Subjects: **Cosmology and Extragalactic Astrophysics (astro-ph.CO)**

Cite as: **arXiv:1001.0013v2 [astro-ph.CO]**

Submission history

From: Tomotsugu Goto [view email]

[v1] Wed, 30 Dec 2009 21:10:33 GMT (227kb)

[v2] Fri, 8 Jan 2010 21:42:51 GMT (232kb)

Download:

- [PostScript](#)
- [PDF](#)
- [Other formats](#)

Current browse context:

[astro-ph.CO](#)

[< prev](#) | [next >](#)

[new](#) | [recent](#) | [1001](#)

Change to browse by:

[astro-ph](#)

References & Citations

- [SLAC-SPIRES HEP](#)
(refers to | cited by)
- [NASA ADS](#)
- [CiteBase](#)

Bookmark (what is this?)



Which authors of this paper are endorsers?

Link back to: [arXiv](#), [form interface](#), [contact](#).