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oscillating neutron stars

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Nonradial superfluid modes in

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outer region of the core but with superfluid intermediate region. We show, that oscillation spectra contain a set of modes whose frequencies can be very sensitive to temperature variations. Fast temporal evolution of the pulsation spectrum in the course of neutron star cooling is also analysed. Comments: 5+ pages, 4 figures, published version. Eqs. (3) and (5) and a few typos are corrected; results unchanged Solar and Stellar Astrophysics (astro-ph.SR); High Subjects: Energy Astrophysical Phenomena (astro-ph.HE); General Relativity and Quantum Cosmology (gr-qc) Journal reference: MNRAS 418 (2011) L54-L58

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For the first time nonradial oscillations of superfluid nonrotating stars are self-

equation of state and realistic density dependent model of critical temperature

consistently studied at finite stellar temperatures. We apply a realistic

of neutron and proton superfluidity. In particular, we discuss three-layer configurations of a star with no neutron superfluidity at the centre and in the

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