



{\it Suzaku} observation of Galactic supernova remnant CTB 37A (G348.5+0.1)

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We present here the results of the observation of CTB 37A obtained with the X-ray Imaging Spectrometer onboard the {\it Suzaku} satellite. The X-ray spectrum of CTB 37A is well fitted by two components, a single-temperature ionization equilibrium component (VMEKAL) with solar abundances, an electron temperature of $kT_e \sim 0.6$ keV, absorbing column density of $N_H \sim 3 \times 10^{22}$ cm^{-2} and a power-law component with photon index of $\Gamma \sim 1.6$. The X-ray spectrum of CTB 37A is characterized by clearly detected K-shell emission lines of Mg, Si, S, and Ar. The plasma with solar abundances supports the idea that the X-ray emission originates from the shocked interstellar material. The ambient gas density, and age of the remnant are estimated to be $\sim 1 f^{-1/2}$ cm^{-3} and $\sim 3 \times 10^4 f^{1/2}$ yr, respectively. The center-filling X-ray emission surrounded by a shell-like radio structure and other X-ray properties indicate that this remnant would be a new member of mixed-morphology supernova remnant class.

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