arXiv.org > astro-ph > arXiv:1107.0826

Search or Article-id

(Help | Advanced search)

All papers



Astrophysics > High Energy Astrophysical Phenomena

Semi-analytical model of cosmic ray electron transport

Alex Ivascenko, Felix Spanier

(Submitted on 5 Jul 2011)

We present a numerical extension to the analytical propagation model introduced in Hein and Spanier (2008) to describe the leptonic population in the galactic disc. The model is used to derive a possible identification of the components that contribute to the leptonic cosmic ray spectrum, as measured by PAMELA, Fermi and HESS, with an emphasis on secondary e+ - e- production in collisions of cosmic ray particles with ambient interstellar medium (ISM). We find that besides secondaries, an additional source symmetric in e+ and e- production is needed to explain both the PAMELA anomaly and the Fermi bump, assuming a power-law primary electron spectrum. Our model also allows us to derive constraints for some properties of the ISM.

Comments: 5 pages, 2 figures

Subjects: High Energy Astrophysical Phenomena (astro-ph.HE)

Journal reference: Astrophys. Space Sci. Trans., 7, 265-269, 2011

DOI: 10.5194/astra-7-265-2011

Cite as: arXiv:1107.0826v1 [astro-ph.HE]

Submission history

From: Alex Ivascenko [view email]

[v1] Tue, 5 Jul 2011 09:20:03 GMT (416kb,D)

Which authors of this paper are endorsers?

Link back to: arXiv, form interface, contact.

Download:

- PDF
- Other formats

Current browse context:

astro-ph.HE

< prev | next >

new | recent | 1107

Change to browse by:

astro-ph

References & Citations

- INSPIRE HEP (refers to | cited by)
- NASA ADS

Bookmark(what is this?)









