

## A Note on Proper Affine Symmetry in Bianchi

### Types $VI_0$ and $VII_0$ Space-Times

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#### Abstract

A study proper affine symmetry in the most general form of the Bianchi types  $VI_0$  and  $VII_0$  space-times is given by using holonomy and decomposability, the rank of the  $6 \times 6$  Riemann matrix and direct integration techniques. It is shown that the very special classes of the above space-times admit proper affine symmetry.

**Keywords:** Affine vector fields, holonomy and decomposability, direct integration technique.

## 1 INTRODUCTION

In this paper we will explore all the possibilities when the Bianchi types  $VI_0$  and  $VII_0$  space-times admit proper affine symmetry. We use holonomy and decomposability, the rank of the  $6 \times 6$  Riemann matrix and direct integration techniques to study proper affine symmetry in the above space-times. Throughout

$M$  represents a four dimensional, connected, Hausdorff space-time manifold with Lorentz metric  $g$  of signature  $(-, +, +, +)$ . The curvature tensor associated with  $g_{ab}$ , through the Levi-Civita connection, is denoted in component form by  $R^a{}_{bcd}$ . The usual covariant, partial and Lie derivatives are denoted by a semicolon, a comma and the symbol  $L$ , respectively. Round and square brackets denote the usual symmetrization and skew-symmetrization, respectively. Here,  $M$  is assumed non-flat in the sense that the curvature tensor does not vanish over any non-empty open subset of  $M$ .

A vector field  $X$  on  $M$  is called an affine vector field if it satisfies

$$X_{a;bc} = R_{abcd} X^d, \quad (1)$$

where  $R_{abcd} = g_{af} R^f{}_{bcd} = g_{af} (\Gamma^f{}_{bd,c} - \Gamma^f{}_{bc,d} + \Gamma^f{}_{ce} \Gamma^e{}_{bd} - \Gamma^f{}_{ed} \Gamma^e{}_{bc})$ . If one decomposes  $X_{a;b}$  on  $M$  into its symmetric and skew-symmetric parts

$$X_{a;b} = \frac{1}{2} H_{ab} + G_{ab}, \quad (H_{ab} (\equiv X_{a;b} + X_{b;a}) = H_{ba}, \quad G_{ab} = -G_{ba}) \quad (2)$$

then equation (1) is equivalent to

$$(i) H_{ab;c} = 0 \quad (ii) G_{ab;c} = R_{abcd} X^d \quad (iii) G_{ab;c} X^c = 0. \quad (3)$$

The proof of the above equation (1) implies (3) or equations (3) implies (1) can be found in [2, 3]. If  $H_{ab} = 2cg_{ab}$ ,  $c \in \mathbb{R}$ , then the vector field  $X$  is called homothetic (and Killing if  $c = 0$ ). The vector field  $X$  is said to be proper affine if it is not homothetic vector field and also  $X$  is said to be proper homothetic vector field if it is not Killing vector field on  $M$  [2]. Define the subspace  $S_p$  of the tangent space  $T_p M$  to  $M$  at  $p$  as those  $k \in T_p M$  satisfying

$$R_{abcd} k^d = 0. \quad (4)$$

## 2 Affine Vector Fields

Suppose that  $M$  is a simple connected space-time. Then the holonomy group of  $M$  is a connected Lie subgroup of the identity component of the Lorentz group and is thus characterized by its subalgebra in the Lorentz algebra. These have been labeled into fifteen types  $R_1 - R_{15}$  [1]. It follows from [2] that the only such space-times which could admit proper affine vector fields are those which admit nowhere zero covariantly constant second order symmetric tensor field  $H_{ab}$ . This forces the holonomy type to be either  $R_2$ ,  $R_3$ ,  $R_4$ ,  $R_6$ ,  $R_7$ ,  $R_8$ ,  $R_{10}$ ,  $R_{11}$  or  $R_{13}$  [2]. A study of the affine vector fields for the above holonomy type can be found in [2]. It follows from [4] that the rank of the  $6 \times 6$  Riemann matrix of the above space-times which have holonomy type  $R_2$ ,  $R_3$ ,  $R_4$ ,  $R_6$ ,  $R_7$ ,

$R_8, R_{10}, R_{11}$  or  $R_{13}$  is at most three. Hence for studying affine vector fields we are interested in those cases when the rank of the  $6 \times 6$  Riemann matrix is less than or equal to three.

### 3 Main Results

Consider Bianchi types  $VI_0$  and  $VII_0$  space-times in usual coordinate system  $(t, x, y, z)$  (labeled by  $(x^0, x^1, x^2, x^3)$ , respectively) with line element [5]

$$ds^2 = -dt^2 + (A f^2(z) + B h^2(z))dx^2 + (A h^2(z) + B f^2(z))dy^2 + 2(A + B)f(z)h(z)dx dy + C(t)dz^2, \tag{5}$$

where  $A, B$  and  $C$  are nowhere zero functions of  $t$  only. For  $f(z) = \cosh z, h(z) = \sinh z$  or  $f(z) = \cos z, h(z) = \sin z$  the above space-time (5) becomes Bianchi type  $VI_0$  or  $VII_0$ , respectively. The above space-time admits three linearly independent Killing vector fields which are

$$\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, -y \frac{\partial}{\partial x} - x \frac{\partial}{\partial y} + \frac{\partial}{\partial z}. \tag{6}$$

The non-zero components of the Riemann tensor are

$$\begin{aligned} R_{0101} &= -\frac{1}{4} \left[ \left( \frac{2A\ddot{A} - \dot{A}^2}{A} \right) f^2(z) + \left( \frac{2B\ddot{B} - \dot{B}^2}{B} \right) h^2(z) \right] = \alpha_1, \\ R_{0102} &= -\frac{1}{4} \left[ \left( \frac{2A\ddot{A} - \dot{A}^2}{A} \right) + \left( \frac{2B\ddot{B} - \dot{B}^2}{B} \right) \right] f(z)h(z) = \alpha_7, \\ R_{0113} &= R_{0223} = -\frac{1}{4ABC} [A^2\dot{B}C + \dot{A}B^2C + 2AB\dot{C}(A+B) - 3ABC(\dot{A} + \dot{B})] f(z)h(z) = \alpha_8, \\ R_{0123} &= -\frac{1}{4} \left[ \left\{ (A+B) \left( \frac{\dot{A}}{A} + \frac{\dot{C}}{C} \right) - 2(\dot{A} + \dot{B}) \right\} h^2(z) + \left\{ (A+B) \left( \frac{\dot{B}}{B} + \frac{\dot{C}}{C} \right) - 2(\dot{A} + \dot{B}) \right\} f^2(z) \right] \\ &= \alpha_9, \quad R_{0202} = -\frac{1}{4} \left[ \left( \frac{2A\ddot{A} - \dot{A}^2}{A} \right) h^2(z) + \left( \frac{2B\ddot{B} - \dot{B}^2}{B} \right) f^2(z) \right] = \alpha_2, \\ R_{0213} &= -\frac{1}{4} \left[ \left\{ (A+B) \left( \frac{\dot{A}}{A} + \frac{\dot{C}}{C} \right) - 2(\dot{A} + \dot{B}) \right\} f^2(z) + \left\{ (A+B) \left( \frac{\dot{B}}{B} + \frac{\dot{C}}{C} \right) - 2(\dot{A} + \dot{B}) \right\} h^2(z) \right] \\ &= \alpha_{10}, \quad R_{0303} = -\frac{1}{4} \left( \frac{2C\ddot{C} - \dot{C}^2}{C} \right) = \alpha_3, \quad R_{0312} = -\frac{1}{4} (A+B) \left( \frac{\dot{A}B - A\dot{B}}{AB} \right) = \alpha_{11}, \\ R_{1212} &= \frac{1}{4ABC} [(A+B)^2 + \dot{A}\dot{B}C] [Af^2(z) + Bh^2(z)] [Ah^2(z) + Bf^2(z)] = \alpha_4, \\ R_{1313} &= \frac{1}{4AB} [\{A^3 - AB(2A + 3B - \dot{A}\dot{C})\} f^2(z) - \{AB(3A + 2B - \dot{B}\dot{C}) - B^3\} h^2(z)] = \alpha_5, \\ R_{1323} &= \frac{1}{4AB} [A^2(A - 5B) + B^2(B - 5A) + AB\dot{C}(\dot{A} + \dot{B})] f(z)h(z) = \alpha_{12}, \\ R_{2323} &= \frac{1}{4AB} [\{A^3 - AB(2A + 3B - \dot{A}\dot{C})\} h^2(z) - \{AB(3A + 2B - \dot{B}\dot{C}) - B^3\} f^2(z)] = \alpha_6. \end{aligned}$$

Writing the curvature tensor with components  $R_{abcd}$  at  $p$  as a  $6 \times 6$  symmetric matrix [6]

$$R_{abcd} = \begin{pmatrix} \alpha_1 & \alpha_7 & 0 & 0 & \alpha_8 & \alpha_9 \\ \alpha_7 & \alpha_2 & 0 & 0 & \alpha_{10} & \alpha_8 \\ 0 & 0 & \alpha_3 & \alpha_{11} & 0 & 0 \\ 0 & 0 & \alpha_{11} & \alpha_4 & 0 & 0 \\ \alpha_8 & \alpha_{10} & 0 & 0 & \alpha_5 & \alpha_{12} \\ \alpha_9 & \alpha_8 & 0 & 0 & \alpha_{12} & \alpha_6 \end{pmatrix}. \quad (7)$$

As mentioned in section 2, the space-times which can admit proper affine vector fields have holonomy type  $R_2, R_3, R_4, R_6, R_7, R_8, R_{10}, R_{11}$  or  $R_{13}$  and the rank of the  $6 \times 6$  Riemann matrix is at most three. Therefore we are only interested in those cases when the rank of the  $6 \times 6$  Riemann matrix is less than or equal to three. Hence there exist following two possibilities when the rank of the  $6 \times 6$  symmetric matrix is less or equal to three which are:

(A) Rank=3,  $\alpha_1 = \alpha_2 = \alpha_3 = \alpha_7 = \alpha_8 = \alpha_9 = \alpha_{10} = \alpha_{11} = 0$ ,  $\alpha_4 \neq 0, \alpha_5 \neq 0, \alpha_6 \neq 0$  and  $\alpha_{12} \neq 0$ .

(B) Rank=1,  $\alpha_4 \neq 0$ , and

$$\alpha_1 = \alpha_2 = \alpha_3 = \alpha_5 = \alpha_6 = \alpha_7 = \alpha_8 = \alpha_9 = \alpha_{10} = \alpha_{11} = \alpha_{12} = 0.$$

We will consider each case in turn.

#### Case (A):

In this case we have  $\alpha_4 \neq 0, \alpha_5 \neq 0, \alpha_6 \neq 0, \alpha_{12} \neq 0$ ,

$\alpha_1 = \alpha_2 = \alpha_3 = \alpha_7 = \alpha_8 = \alpha_9 = \alpha_{10} = \alpha_{11} = 0$ , the rank of the  $6 \times 6$  Riemann matrix is three and there exists a unique (up to a multiple) no where zero time like vector field  $t_a = t_{,a}$  solution of equation (4) and  $t_{a,b} \neq 0$ . From the above constraints we have  $B(t) = dA(t)$ ,  $C(t) = eA(t)$  and  $A(t) = (at + b)^2$ , where  $a, b, d, e \in R (d, e > 0)$ . The line element in this case takes the form

$$ds^2 = -dt^2 + (at + b)^2 \left[ \begin{aligned} & (f^2(z) + eh^2(z))dx^2 + (dh^2(z) + f^2(z))dy^2 + \\ & 2(1+d)f(z)h(z)dx dy + e dz^2 \end{aligned} \right]. \quad (8)$$

Substituting the above information into the affine equations one find that

$$X^0 = c_4 t + c_5, X^1 = -c_1 y + c_2, X^2 = -c_1 x + c_3, X^3 = c_1, \quad (9)$$

where  $c_1, c_2, c_3, c_4, c_5 \in R$ . One can write the above equation (9) subtracting Killing vector fields as

$$X = (c_4 t + c_5, 0, 0, 0). \quad (10)$$

Clearly in this case one can easily see that the above space-times (8) admit proper affine symmetry. It is important to note that the constants  $a$  and  $b$  can not be zero simultaneously.

**Case (B):**

In this case we have

$\alpha_1 = \alpha_2 = \alpha_3 = \alpha_5 = \alpha_6 = \alpha_7 = \alpha_8 = \alpha_9 = \alpha_{10} = \alpha_{11} = \alpha_{12} = 0$ ,  $\alpha_4 \neq 0$  and the rank of the  $6 \times 6$  Riemann matrix is one. Here, there exist two linear independent solutions  $t_a = t_{,a}$  and  $z_a = z_{,a}$  of equation (4). The vector field  $t_a$  is covariantly constant whereas  $z_a$  is not covariantly constant. From the above constraints we have  $A(t) = B(t) = C(t) = (t + q)^2$ , where  $q \in R$ . The line element takes the form

$$ds^2 = -dt^2 + (t + q)^2 \left[ \frac{(f^2(z) + h^2(z))dx^2 + (h^2(z) + f^2(z))dy^2 + 4f(z)h(z)dxdy + dz^2}{4f(z)h(z)dxdy + dz^2} \right]. \tag{11}$$

Affine vector fields in this case

$$X^0 = c_4t + c_5z + c_6, X^1 = -c_1y + c_2, X^2 = -c_1x + c_3, X^3 = c_1, \tag{12}$$

where  $c_1, c_2, c_3, c_4, c_5, c_6 \in R$ . One can write the above equation (12) subtracting Killing vector fields as

$$X = (c_4t + c_5z + c_6, 0, 0, 0). \tag{13}$$

In this case clearly the above space-times (11) admit proper affine symmetry.

**SUMMARY**

In this paper an attempt is made to explore all the possibilities when the Bianchi types  $VI_0$  and  $VII_0$  space-times admit proper affine vector fields. A different approach is adopted to study proper affine vector fields of the above space-times by using holonomy and decomposability, the rank of the  $6 \times 6$  Riemann matrix and direct integration techniques. From the above study we obtain the following results:

- (i) We obtain the space-time (8) that admits proper affine vector fields when the rank of the  $6 \times 6$  Riemann matrix is three and there exists a unique nowhere zero independent timelike vector field, which is the solution of equation (4) and is not covariantly constant (for details see case A).
- (ii) The space-time (11) is obtained, which admits proper affine vector fields (see case B) when the rank of the  $6 \times 6$  Riemann matrix is one and there exist two independent solutions of equation (4) but only one independent covariantly constant vector field.

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**Received: March, 2010**